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1 **Automated collection and categorisation of STM images and STS** 2 **spectra with and without Machine learning**

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7 **Abstract**

8 **Background:** : Atomic resolution scanning probe microscopy, and in particular scanning tun-
9 nelling microscopy (STM) allows for high spatial resolution imaging, and also spectroscopic anal-
10 ysis of small organic molecules. However, preparation, and characterisation of the probe apex *in*
11 *situ* by a human operator is one of the major barriers to high throughput experimentation, and to re-
12 producibility between experiments. Characterisation of the probe apex is usually accomplished via
13 assessment of the imaging quality on the target molecule, and also the characteristics of the scan-
14 ning tunnelling spectra (STS) on clean metal surfaces. Critically for spectroscopic experiments,
15 assessment of the spatial resolution of the image is not sufficient to ensure a high quality tip for
16 spectroscopic measurement. The ability to automate this process is a key aim in development of
17 high resolution scanning probe materials characterisation.

18 **Results:** In this paper, we assess the feasibility of automating the assessment of imaging quality,
19 and spectroscopic tip quality, via both machine learning (ML) and deterministic methods (DM) us-
20 ing a prototypical Tin Phthalocyanine (SnPc) on Au(111) system at 4.7 K. We find that both ML
21 and DM are able to classify images and spectra with high accuracy, with only a small amount of
22 prior surface knowledge. We highlight the practical advantage of DM not requiring large training
23 datasets to implement on new systems and demonstrate a proof-of-principle automated experiment
24 that is able to repeatedly prepare the tip, identify molecules of interest and perform site specific

25 STS experiments using DM, in order to produce large numbers of spectra with different tips suit-
26 able for statistical analysis.

27 **Conclusion:** Deterministic methods can be easily implemented to classify the imaging and spec-
28 troscopic quality of a STM tip for the purposes of high resolution STM and STS on small organic
29 molecules. Via automated classification of the tip state, we demonstrate an automated experiment
30 that can collect high number of spectra on multiple molecules without human intervention. The
31 technique can be easily extended to most metal-adsorbate systems, and is promising for the devel-
32 opment of automated, high-throughput, STM characterisation of small adsorbate systems.

33 **Keywords**

34 STM; STS; Machine Learning; Automated; spectroscopy;

35 **Introduction**

36 STS extends the capability of STM beyond topographic imaging, allowing for the direct measure-
37 ment of the electronic properties of surfaces and molecules with atomic precision. This opens up
38 the ability to map the local density of states (LDOS) of a sample with high spatial resolution [1-3].
39 Peaks within a map of the LDOS correspond to increases in conductance at specific bias values, re-
40 vealing the energy levels of key features (e.g. molecular orbitals in the case of molecular samples)
41 within the material.

42 As for STM imaging, the sharpness and overall tip shape is crucial in optimising the spatial res-
43 olution of STS measurements; sharp tips result in localised tunnelling through a single position,
44 whereas blunt or misshaped tips cause averaging of contributions over larger areas, reducing the
45 spatial resolution and potentially blending the electronic features between different sites. However,
46 even for tips with high spatial resolution, it is known that different tip structures and probe termina-
47 tions are known to influence these results [4-7].

48 STS measurements are the result of an integration over the available density of states (DOS) in
49 both the tip and the sample, with the current measured therefore being proportional to the con-
50 volution of two. To isolate the DOS of the sample, it is crucial that the tip has a nominally “flat”

51 DOS, which is typically achieved by using a purely metallic tip. However, most tips do not demon-
52 strate a perfectly flat LDOS as they have a complex electronic structure governed by the geometry
53 of the metallic cluster at the tip apex [8-12]. Non-metallic contaminants can also strongly perturb
54 the electronic structure of the tip.

55 Methods of optimising the probe state for ideal STS are slow and laborious, involving indentation
56 into a metal surface and bias pulses applied to the tip, manually checking spectra and imaging af-
57 ter each probe shaping attempt. The automation of this process could result in a more rapid and
58 reproducible method for performing spectroscopy measurements.

59 To classify the state of the probe for STS experiments, spectra are usually taken over bare areas of
60 a metallic substrate. On coinage metal surfaces, these $\frac{dI}{dV}$ spectra typically exhibit a characteristic
61 feature corresponding to the surface state, which appears as a step function around a specific bias
62 value, which for the Au(111) surface appears at around -0.48 V [13,14].

63 One notable attempt to automate this classification using machine learning (ML) was carried out by
64 Wang *et al.* [15]. This work aimed to classify the state of a STM tip based on STS measurements
65 of the Au(111) surface. Using a total of 1789 archived $\frac{dI}{dV}$ spectra, a ML model was trained which
66 aimed to classify new spectra into one of five categories, based on the similarity of the spectra to an
67 idealised surface state.

68 This schema achieved final precision in classification of 84%, and a recall of 74%. Similarly to
69 image classification in scanning probe microscopy (SPM) [16-18], the availability of such a large
70 amount of data for training is usually very low, making ML based classifiers troublesome to train.
71 In addition to the lack of data, ML models require careful labelling and a high level of knowledge
72 from the labeller to be able to train such a model. Switching to a new substrate system is likely to
73 require retraining of the model, and furthermore, even after a successful training, it is still often un-
74 clear what the model is learning from the input data, a problem which leads to these models being
75 referred to as a ‘black box’. Because of these limitations, there is a strong case to develop methods
76 which do not rely on ML, circumventing these drawbacks whilst still able to make precise classifi-
77 cations of the tip quality for use in automation.

78 In the following work, we use the prototypical system of Tin Phthalocyanine (SnPc) on Au(111)
79 to investigate the feasibility of a DM automated classifier, and compare it to ML methods. This
80 molecular system has the advantage that the SnPc adsorbs on the surface in two distinct configu-
81 rations, one with the tin atom facing up (SnUp), and the other with the tin facing down (SnDown)
82 (Figure 1), providing a variety of molecular configurations to challenge the automated molecular
83 identification.

84 In addition to the classification of the probe quality based on the surface state, we use a DM based
85 cross-correlation (CC) feature finding method [19] in order to also assess the imaging quality of the
86 tip, and also automatically locate various molecules on the surface. Combining these methods, it is
87 possible to conduct a fully automated experiment, where a large number of STS measurements can
88 be obtained over various molecules automatically, with optimised tips and the quality of the spectra
89 and image assessed automatically.

90 Methods

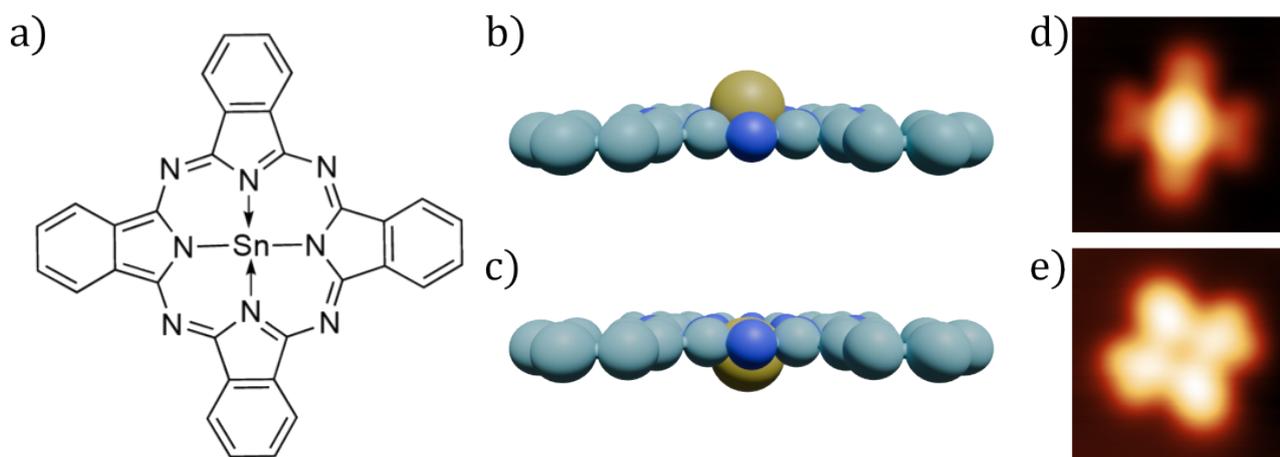


Figure 1: a) Structure of tin phthalocyanine (SnPc). Side-on view of SnPc, illustrating its non-planar nature in the b) SnUp and c) SnDown configurations. d) and e) show constant current STM images of the SnUp and SnDown configurations respectively, taken at 5 K, -100 mV, 50 pA.

91 **Experimental details**

92 We used a third generation commercial low temperature (LT) STM NC-AFM instrument (Scienta
93 Omicron GmbH) operating in UHV (base pressure $\leq 5 \times 10^{-11}$ mbar) cooled to 5 K. Gold and silver
94 crystals (spl.eu) were prepared via repeated sputter-anneal cycles, sputtering under an Argon pres-
95 sure of $\sim 5 \times 10^{-5}$ mbar, with a beam energy of 1.5 kV for 30 minutes, measuring a drain current of
96 $\sim 7.0 \mu\text{A}$, before annealing at 500 °C for 30 minutes and then placed into the scan head for imag-
97 ing. Platinum iridium STM tips were used throughout this work and were prepared by standard
98 STM methods (voltage pulses, controlled contacts with the sample) until good atomic resolution
99 was obtained in STM feedback.

100 SnPc was deposited onto the Au(111) surface using a home-built evaporator, where the powdered
101 source material is contained within a glass crucible using glass wool, around which a coil of tan-
102 talum wire is wound, providing a source of heat for the crucible. The target temperature for SnPc
103 deposition was 360 °C, once reached, the cryostat shields were opened for 1 hour, before closing
104 and checking the coverage in STM. Once deposited, the sample was cold annealed to room tem-
105 perature, which has the effect of driving the molecules preferentially to the “elbow” sites of the
106 herringbone structure.

107 An STS spectrum (differential conductance) can be obtained in practice using one of two methods.
108 Both begin by positioning the STM tip at a desired lateral position on the surface, whilst scanning
109 in STM feedback. At this point, the feedback loop is disabled, keeping the tip-sample distance con-
110 stant throughout the spectroscopy measurement. The voltage is then swept through a range of val-
111 ues whilst measuring the current, which is obtained as a function of the varying voltage, $I(V)$. This
112 curve can then be differentiated with respect to the voltage to obtain the differential conductance,
113 $\frac{dI}{dV}$, spectra.

114 Alternatively, the derivative signal, $\frac{dI}{dV}$, as a function of voltage, can be directly measured using
115 the lock-in technique. In this method, an AC signal is generated by applying a small modulation
116 voltage, $V_M \cos(\omega t)$, to the bias. Due to this modulation, the measured current is expressed as

117
$$I(t) = f(V + V_M \cos(\omega t)) \tag{1}$$

118 Where V_M is the modulation amplitude and ω is the frequency. Applying this modulation around
119 a central voltage, creates a corresponding modulation in the measured tunnel current, with an am-
120 plitude proportional to the gradient of the $\frac{dI}{dV}$ curve at that bias. Therefore, once the tip is in posi-
121 tion, the bias can be swept through a range, whilst applying the modulation. The resultant current
122 can then be detected by a lock-in amplifier, where its amplitude for small values of V_M is propor-
123 tional to $\frac{dI}{dV}$, therefore directly measuring the differential conductance of the sample. Throughout
124 the work presented here, the conductance was measured directly using the lock-in technique.

125 **Results**

126 To create the ML based classification models needed for this work, a large amount of data was
127 needed in the form of STS spectra taken with a variety of different tip shapes and configurations.
128 The process of the dataset generation was performed in a manner similar to that described in
129 Barker et al. [19], with some minor alterations, as described below.

130 The process of the automated dataset generation is shown in Figure 2. One addition to this method
131 compared to the automated data gathering method described in Barker et al. [19] is the addition
132 of $I(z)$ classifications prior to performing imaging to ensure a tunnelling junction. This acts as a
133 rapid “pre-filtering” step, eliminating tips that do not show a stable tunnelling junction (and hence
134 are not suitable for STS) without the need to perform a complete image to characterise the tip. The
135 classification of the state of the probe based on imaging is performed, via the CC method, as a key
136 feature of a “good” tip for STS is also the sharpness of the probe, in order to ensure high spatial
137 resolution in the acquired data. Further details on the CC and $I(z)$ classification as implemented for
138 the SnPc models is provided in the online supporting information.

139 Once the imaging classification is complete, the $I(z)$ classification is performed again to check that

140 a tip change did not occur during the scan. The obtained topograph is then analysed to find both
141 a large area of clean metal substrate, over which the I(V) spectra can be obtained, and to find the
142 location of the molecules in various configurations, over which additional I(V) spectra are taken.
143 After completing this data gathering step, the tip moves away from the imaging area for a tip prepa-
144 ration event, in order to change the apex substantially before repeating the entire process to col-
145 lect another dataset with a different tip. Throughout, if the tip is classified as “bad” in either of the
146 I(z) classifications or the CC based imaged classification, the script moves onto a tip preparation
147 event. If the tip fails in being classified as “good” more than a set number of times in a row, the
148 tip is moved away macroscopically, under the assumption that the current area of the surface is not
149 suitable for classification; this typically occurs due to the area being damaged from prior tip prepa-
150 ration, or the absence of an SnUp molecule in the frame which is used in the CC classification of
151 the image.

152 The CC classification is carried out as described in Barker et al. ([19]), with the reference image
153 used being a cropped image of an SnUp molecule as is shown in Figure 1d. SnUp molecules were
154 chosen for classification as in this configuration, the Sn atom in the molecule presents a higher as-
155 pect ratio than in the SnDown configuration, and so is more sensitive to the sharpness of the tip.
156 Using this method with a threshold of > 0.98 , the model was able to reliably generate and identify
157 sharp tips.

158 **Dataset summary**

159 Using the data generation method described above, we obtained a total of 2604 individual spec-
160 tra on the bare Au(111) surface to use for our classification models. In order to use this data
161 for training and evaluation of models, the dataset required labelling. We note that labelling of the
162 dataset is non-trivial, as for ML models the model can only attempt to ‘learn’ to evaluate spectra
163 based on the ‘ground truth’ provided by the labelling.

164 Labelling was carried out using a similar process to Barker et al. [19]. A custom Python script was
165 written with a graphical interface. The script would show each spectra individually, with a choice
166 of four labels depending on the visibility of the surface state (SS): SS “good”, SS “step visible”, SS

167 “peak visible” and SS “not visible”. When classifying the data, the region around the surface state
168 step was focused on, with the “good” label being attributed to a spectra where the step was clearly
169 visible at the correct position, with few features before and after the step.

170 Whilst it was clear which data fell into each category, we note that even the data with the most
171 visible surface state contained a background slope, as seen in Figure 3c. This slope most likely
172 arises from a macroscopic property of the tip that was not removed by the moderate tip treatment
173 steps used during the automated process, which did not include aggressive treatment such as high
174 current field emission. Our data were therefore classified with this trend in mind, with the final
175 “good” classifications often containing a trend in the < -0.5 V region. The SS “step visible” label
176 was given to spectra whose curves show the step in the correct position, but where a general trend
177 (slope) was also visible throughout the data. The SS “peak visible” label applied to spectra where
178 there was an apparent feature at the correct position for the step, but not necessarily a step, and the
179 final SS “not visible” label was given to spectra where no feature resembling the surface state could
180 be observed. A representative sampling of spectra from each labelling category are shown in Fig-
181 ures 3 and 4, with all spectra obtained shown in the online supplementary information S2.

182 For the final classification, the SS “good” and SS “step visible” categories were combined into a
183 single “good” category, and the SS “not visible” and SS “peak visible” were combined into a “bad”
184 category. This was done to allow for a simple binary classification on the basis that further dis-
185 tinction between the classes is unlikely to improve the final model, and would greatly increase the
186 complexity of the problem.

187 Table 1 shows the number of spectra in each category after the labelling step. For ML training, the
188 data were split into training, validation and test sets at a ratio of 70:10:20. This left 1823 spectra for
189 training, 260 for validation and 521 for final testing.

Table 1: Number of spectra in each category.

Labels	Count
SS “good”	384
SS Step Visible	482
SS Peak Visible	1169
SS Not Visible	569
<hr/>	
Binary “good”	866
Binary “bad”	1738

190 **Classification Methods**

191 **Machine Learning Classifier**

192 With the labelling completed, it was possible to train a series 1D convolutional neural networks
193 (CNNs). In total, 72 models were trained, varying the number of convolutional layers between 1–3,
194 the number of dense training layers between 1 – 3, the number of kernels in the first convolutional
195 layer (32 and 64 kernels were used, doubling in successive layers), kernels of sizes 3×3 or 5×5 , and
196 dropout layers with rates of either 0.3 or 0.5, including all combinations of these. The training was
197 carried out on the training dataset containing 1823 spectra, validating the model after each epoch
198 on the validation set of 260.

199 After training, each of these models were evaluated on a test set of 521 spectra, with their final ac-
200 curacies, precisions and recalls compared. We note the recall is defined as the percentage of all
201 data labelled as “good”, which is then also classified as “good”. This metric therefore places more
202 weight on false negatives than the precision metric and is also not as largely skewed by imbalanced
203 datasets as the accuracy metric.

204 The model which achieved the best balance between the three metrics was one which contained
205 2 convolutional layers starting with 5×5 kernels, 32 in the first layer and 64 in the second, a sin-
206 gle dense training layer, followed by a dropout rate of 0.3. The architecture of this model is shown
207 schematically in Figure S2. This achieved an overall accuracy of 86%, a precision of 85% and a
208 recall of 70%.

209 **Deterministic Classifier**

210 For the deterministic classifier, we required a method which is able to adapt to the entire dataset
211 with a clear set of rules, outputting a metric describing how close any individual spectra is to an
212 idealised surface state spectra. To this end, we implemented a simple model to calculate the dif-
213 ference between the surface state step at -0.48 V with a perfect step function, both normalised be-
214 tween 0 – 1.

215 In principle, for a “ideal” metallic tip, the spectra would appear completely flat on either side of
216 the step function. However, as noted above, the majority of the data we acquired were not com-
217 pletely flat and showed a noticeable slope even when the SS was clearly visible. Therefore, in or-
218 der to make a comparison between these tips and the ideal step function, additional processing is
219 needed.

220 Firstly, the spectra were cropped to remove features outside of the categorisation window, which
221 for this dataset was the bias range -0.55 V \rightarrow 0.5 V. From here, any general trend/slope visible in
222 the data needs to be found and subtracted from the step. Commonly in our data, it seems that the
223 trend is a linear offset in the $\frac{dI}{dV}$, and hence a linear function after the step can be fit to the data, and
224 then subtracted from the original spectrum. The specific location of the turning point of the each
225 spectrum is obtained by finding the minimum of the differential of the curve within a small range
226 around -0.48 V, and the step is assumed to be contained within the categorisation range, follow-
227 ing this determined turning point. From this, a linear function is fit to the window, an example of
228 which is shown for both a “good” and “bad” classified tip in Figure 5a) and c) respectively.

229 Once the linear trend is found, it is subtracted from the original spectrum, the result of which is
230 shown in Figure 5b and d. For a “good” spectrum, the resultant curve should appear roughly as a
231 step function, and so by direct comparison to a perfect step function, starting at the turning point
232 found earlier, a deterministic classification measure can be made. The specific metric output as the
233 difference between these two curves is the root mean squared (RMS) error between the two, which
234 is described by Equation 2

$$RMS = \sqrt{\sum_{i=1}^n (\hat{y}_i - y_i)^2} \quad (2)$$

Here, \hat{y} are the perfect step function data points, and y are the spectra for classification. To evaluate the optimal thresholds for classification, a stacked histogram was plotted, showing the spread of the RMS in each category. This histogram is shown in Figure 6. From this, the final threshold chosen < 0.25 , with all spectra resulting in a value within this range classified as “good”. Using this method, the deterministic model was able to achieve an overall accuracy of 82%, a precision of 86% and a recall of 53% when evaluated on the same test set used for the ML model.

Results and Discussion

Both the deterministic and ML based models were tested on the same isolated test set of 521 spectra, with the final results as given in Table 2. Both models achieve very similar accuracies and precisions, however the recall for the deterministic model is significantly lower than in the ML model. In practise, this lower recall would mean that more tips which a human may classify as “good” would be misclassified as “bad”, slowing down the overall tip preparation process. However, since the precisions of both the ML and deterministic models are very similar, the probability of an automated tip preparation script exiting with a “bad” tip would be roughly the same using either model. Since both models show comparable results in the precision of the final classification, the main advantage to using the deterministic model over ML is that the classifier requires much less labelled data for its creation, and hence is easier to apply to a new system. Our ML accuracies are consistent with the prior work undertaken by Wang *et al.* [15]. Their highest ML based classifier which was able to achieve a precision of 84% and a recall of 74%, whereas our DM based results are substantially better than the DM approach they trialled which used correlation based metrics, and only achieved a final precision of 41% and a recall of 53% (no accuracies were given for this work).

258 It should be noted, however, that Wang *et al.* attempted to make a classification between five differ-
259 ent labels of spectra, whereas our dataset was split into a binary “good” or “bad” before training. In
260 general, binary classifiers are expected to achieve higher accuracies as the differences between the
261 two categories are less subtle.

Table 2: Table showing the accuracy, precision and recall obtained using deterministic and ML models to classify probe tips based on spectroscopy measurements.

	Deterministic	ML
Accuracy	82%	86%
Precision	86%	85%
Recall	53%	70%

262 **Automated Experiment Discussion**

263 In addition to automatically classifying the tip quality via STS on the Au(111) substrate, the script
264 automatically located each SnPc molecule on the Au(111) surface, identified the different config-
265 urations of the molecule, and carried out lock-in $\frac{dI}{dV}$ measurements over the centre of each. In this
266 section, we will discuss the STS data taken on the molecules, discuss the impact on the STS data
267 quality due to the quality of the tip and highlight the advantages of automated assessment of tip
268 quality and statistical categorisation of the data in STS.

269 **Molecule location and identification**

270 Once a series of Au(111) surface spectra had been obtained for use in the classifier training, the
271 script continued on to obtain STS measurements over the centre of each located SnPc molecule,
272 whilst also distinguishing between the two configurations (SnUp and SnDown) prior to measure-
273 ment. The identification of each molecular configuration was determined using the CC method
274 with two separate reference images as shown in Figure 1d-e.

275 For the final distinctions on the Au(111) surface, the CCR thresholds used for the SnUp and
276 SnDown molecules were 0.983 and 0.980 respectively. Using these thresholds on a small test set

277 of 13 images, the script was able to locate the positions of SnUp molecules with a 100% accuracy
278 and precision, whereas on the SnDown the accuracy achieved was 95% with a precision of 96%.
279 Once located, lock-in $\frac{dI}{dV}$ curves were obtained over the central atom of each molecule located, us-
280 ing the same -1.5 V to 1.5 V range, as for the bare surface.

281 **Note on SnPc Switching Instability**

282 The adsorption of SnPc on coinage metal substrates is well studied, and the molecule is known to
283 undergo an irreversible switch from the SnUp to the SnDown state on the Ag(111) surface [20] un-
284 der hole injection. This is usually carried out intentionally by positioning the tip over the centre of
285 an SnUp molecule, and applying a bias pulse via the tip of < -1.9 V. On injection, the Sn atom
286 within the molecule is transiently oxidised to Sn^{3+} , which favours a new position closer to the sur-
287 face, where the atom binds to the Ag(111), at which point charge transfer from the substrate to the
288 molecule will return it to its neutral state [20].

289 Whilst carrying out bias spectroscopy over these SnPc molecules, it was found that a switch could
290 occasionally be induced, even if the bias range used did not reach -1.9 V. With moderate negative
291 bias (e.g., -1.5 V) a switch would commonly occur and even with parameter adjustments to reduce
292 the probability of switching (i.e. reduced integration times), there was still a chance that the switch
293 would be induced, as can be seen in Figure S3.

294 In the automated experiment, an image would be taken, the molecules located based on this image,
295 and then spectra would be obtained, saving each spectra with a label indicating which configura-
296 tion of molecule the spectra was taken over. Unfortunately, due to this switching occurring over
297 specifically the SnUp molecules, spectra labelled SnUp had the potential to be unreliably labelled.
298 Additionally, it was observed that these switches could occur at the start of the measurement in the
299 initial setting of the bias, or during the bias sweep itself, meaning identifying if a switch had oc-
300 curred could not be reliably inferred from simple analysis of the STS spectra. For this reason, the
301 results shown in the next section will only consider spectra taken over the SnDown configuration,
302 as the labelling of these was reliable.

303 **Results**

304 Using our STS spectra surface state classification method as described previously, we were able
305 to process the entire dataset collected, and categorise the spectra taken on molecules as being ac-
306 quired with either a “good” tip, or with a “bad” tip.

307 Throughout the data gathering, a total of 86 images (and so probe tips) passed the $I(z)$ and CCR
308 pre-classification steps and were used to obtain molecular STS measurements. Of these 86 tips,
309 30 were classified as “good” and 56 as “bad”, based on the analysis of the final STS spectra on
310 Au(111). These 30 “good” tips were used to obtain spectra over a total of 49 SnDown molecules.
311 The mean of these curves is shown in Figure 7, where it can be clearly seen that there is a peak at
312 roughly 0.8 V which is not present in the bare surface spectra seen in Figure 3c.

313 Previous STS data of SnDown molecules show a clear increase in conductance at both -0.85 V and
314 0.75 V when imaging at a setpoint of 50 pA [21]. These peaks in conductance correspond to the
315 lowest unoccupied molecular orbital (LUMO) and highest occupied molecular orbital (HOMO)
316 respectively. The cited work, however, was carried out on the Ag(111) surface, as opposed to
317 Au(111) used here, which could explain the slight shift in the position of the HOMO from the lit-
318 erature value of 0.75 V to our consistent measurement of roughly 0.8 V. In addition, the work also
319 suggests a current dependence on the position of the HOMO, which could be a contributing factor
320 to the difference.

321 When comparing the “good” $\frac{dI}{dV}$ spectra taken over the molecules (Figure 7) to the binary “good”
322 bare surface spectra (Figure 3c), the surface state and general increase in conductance at biases be-
323 low -0.5 V can be seen in both. However, unlike the HOMO, which is clearly visible in the molec-
324 ular $\frac{dI}{dV}$, the LUMO is not visible at the expected bias value of -0.85 V. This is possibly due to
325 the peak being obscured by the shoulder in the negative portion of the spectra. Comparing the re-
326 gion between -1.5 and -0.5 V in both spectra, it can be seen that the mean curve for the molecular
327 spectra contains an additional feature which is not present in the mean bare surface state spectrum.
328 The features contained within this could contain the LUMO, but this is difficult to ascertain without
329 completely deconvolving the tip and sample LDOS.

330 Figure 8 shows the mean and standard deviation of the $\frac{dI}{dV}$ spectra taken over SnDown molecules
331 taken with a “bad” tip. By comparing this to Figure 7 it is clear that the HOMO peak at 0.8 V is
332 much less prominent. In addition to this, the features throughout the spectra have become less evi-
333 dent. This clear difference between the molecular STS taken with a “good” and “bad” tip, with the
334 former showing expected features, reinforces that the tip state classification was successful in pro-
335 ducing higher quality spectra, and highlights the importance of appropriately characterizing the tip
336 state before STS experiments.

337 A clear advantage to performing automated experiments with a large number of different tips and
338 over a large number of molecules in different surface positions, is that statistically, variations in the
339 spectra due to the changes in the tip or small changes in the molecular adsorption, will be averaged
340 out, and better approach those from ensemble techniques. As can be seen in both Figures 7 and 8,
341 there is a substantial variation in the individual spectra around the mean curve. This is most likely
342 due to variations in the quality of the tip, or slight differences in the molecule itself. However, with
343 a large enough aggregate of different tips, and taking STS measurements over different molecules,
344 when averaged, these small variations should be dominated by the consistent features present in all
345 the data. This can be seen particularly well in Figure 7, where some of the individual molecular
346 spectra (grey curves), which here were all taken with tips classified as “good”, show a featureless
347 region around the HOMO, whilst others clearly show a strong peak. With individual spectra, it is
348 possible that specific features in the $\frac{dI}{dV}$ could remain unobserved due to spurious problems with the
349 tip.

350 For a human operator, taking a large number of spectra, with different, yet still “good” tips, on dif-
351 ferent instances of the same molecule is extremely time consuming. However, with the entire pro-
352 cess being automated, this can be carried out very simply, and without any need for constant moni-
353 toring.

354 We note that while we collected data using both types of tip in order to highlight the differences
355 in quality, in a real use case data collection would be improved by using the proposed method of
356 classifying the tip based on a $\frac{dI}{dV}$ spectrum taken over the bare surface, such that the script would

357 only take molecular spectra using probes which have been classified as “good” based on the surface
358 state spectra. An example flow diagram, with data taken from a generation run where the surface
359 state spectrum was classified as “good” is shown in Figure 9.

360 **Conclusion**

361 We have shown that it is possible to perform a fully automated experiment, carrying out STS mea-
362 surements over targeted areas of specific organic molecules, including the ability to modify and
363 characterise the state of the tip, by both analysis of its spectroscopic characteristics, and imaging
364 quality, without the use of machine learning. This enables the ability to obtain a large number of
365 spectra over various features on a surface, with a variety of characterised tip states, without the
366 need for an operator to be present, and to perform statistical analysis of the spectroscopic data, via
367 the automated labelling of the state, and location of the spectrum.

368 Importantly, the ability to carry this out without machine learning means that this method can be
369 easily adapted to different adsorbate/substrate systems without the need for extensive data collec-
370 tion to train ML models. This methodology can aid in the rapid characterisation of new materials
371 via automated probing of different features in a system, taking numerous measurements over differ-
372 ent areas, only requiring an operator once the experiment is complete, to process the resultant data
373 for analysis.

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378 BOND.

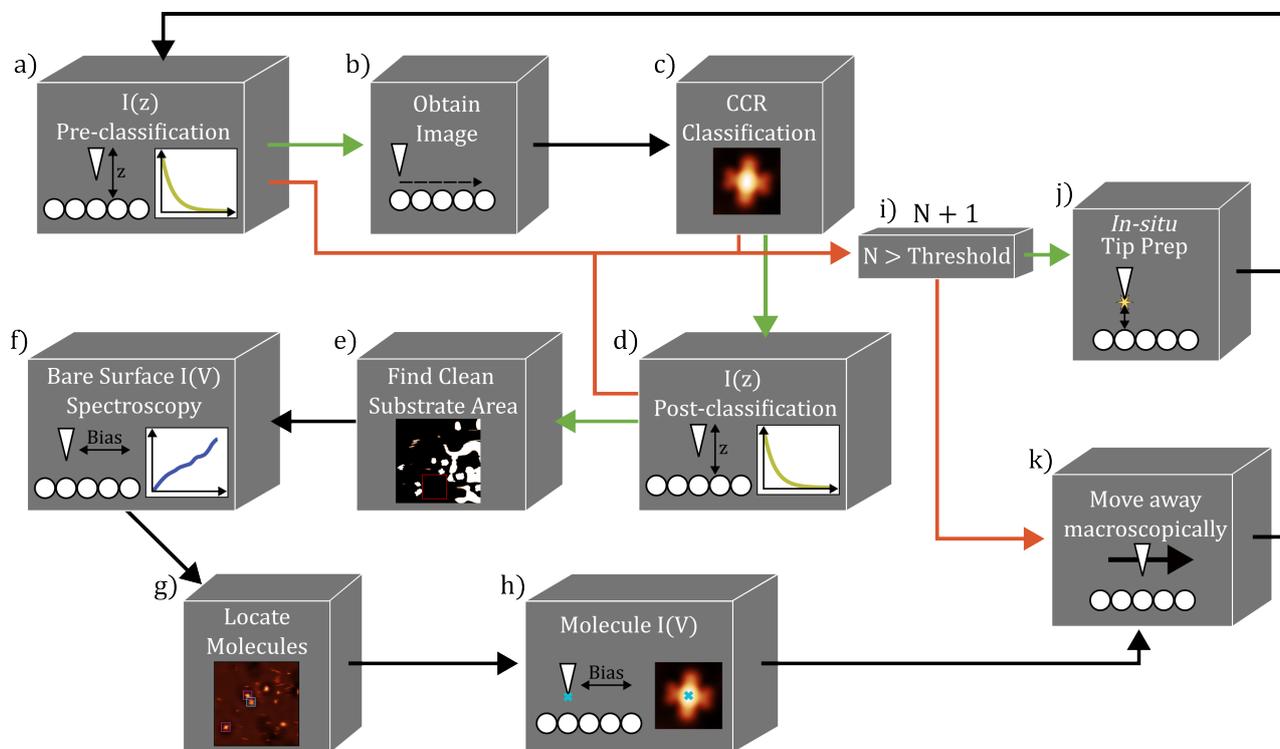


Figure 2: Schematic for the automated data gathering script. The script starts by taking an initial $I(z)$ spectra, a), which is classified based on its exponential dependence. If the $I(z)$ is classified as “good”, the script then obtains a scan of a specific area, b), followed by a CC based classification, c). If the CC image classification determines the tip to be “good”, the script moves onto another $I(z)$ classification, d), followed by an analysis step to find a clean substrate area, e). Using the area found in e), the script obtains 15 $I(V)$ spectra over different positions, f). The script then locates the different configuration of molecules present in the scan, g), before obtaining $I(V)$ spectra over the centre positions of each molecule, h). In the classification steps a), c) and d), if the tip is classified to be “bad”, the script will move on to either an *in situ* tip preparation step, j), or if the number of shaping attempts without a “good” tip has exceeded a pre-determined threshold, i), the tip is moved away macroscopically, k), under the assumption that the area is not suitable for classification. Throughout the schematic, green arrows show positive classifications and red arrows indicate negative classifications.

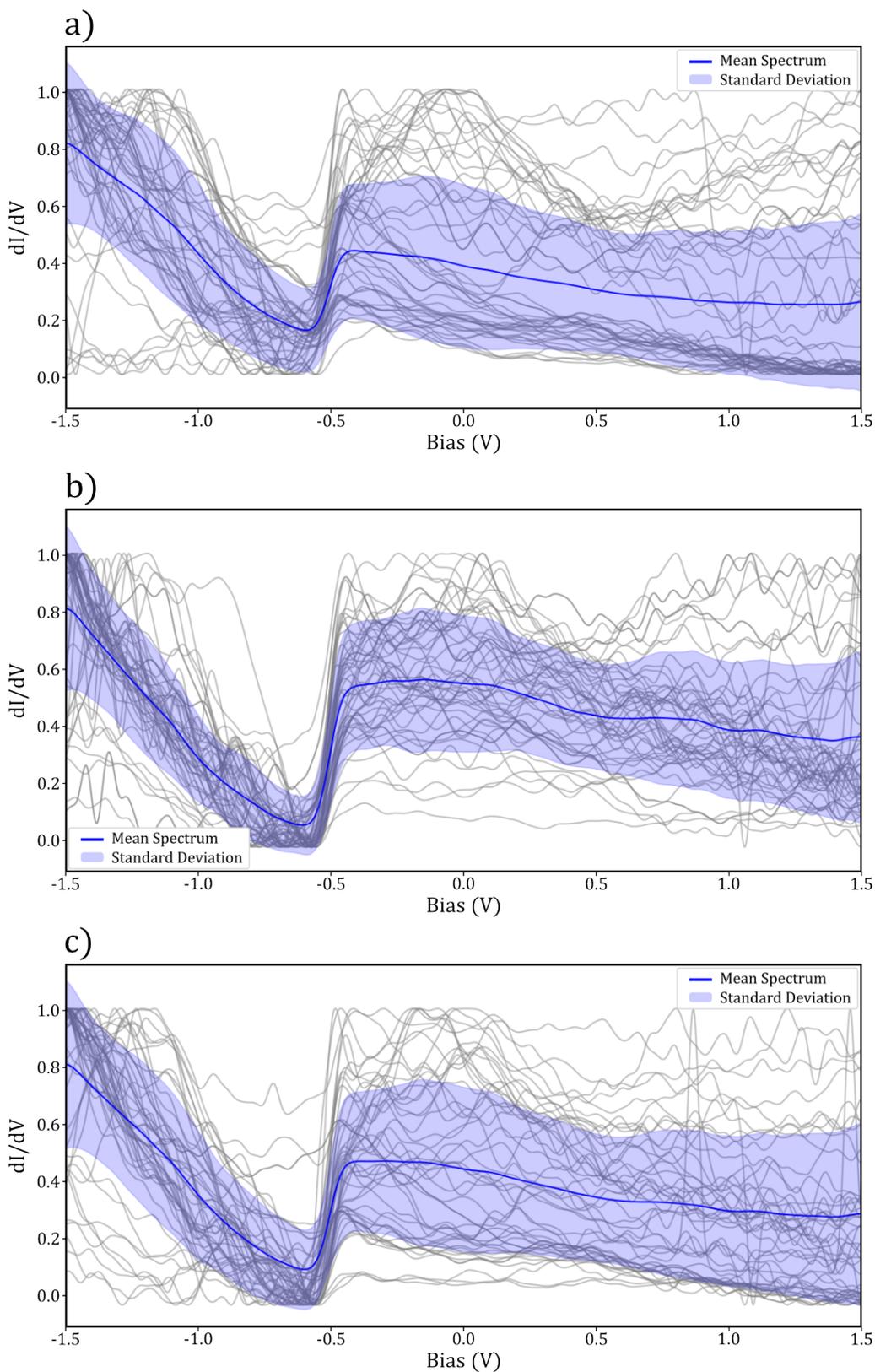


Figure 3: Samples of 50 normalised spectra taken over the clean Au(111) surface, mean (blue line) and standard deviation (shaded area) for a) Surface state step visible, b) surface state “good” and c) binary “good” labels.

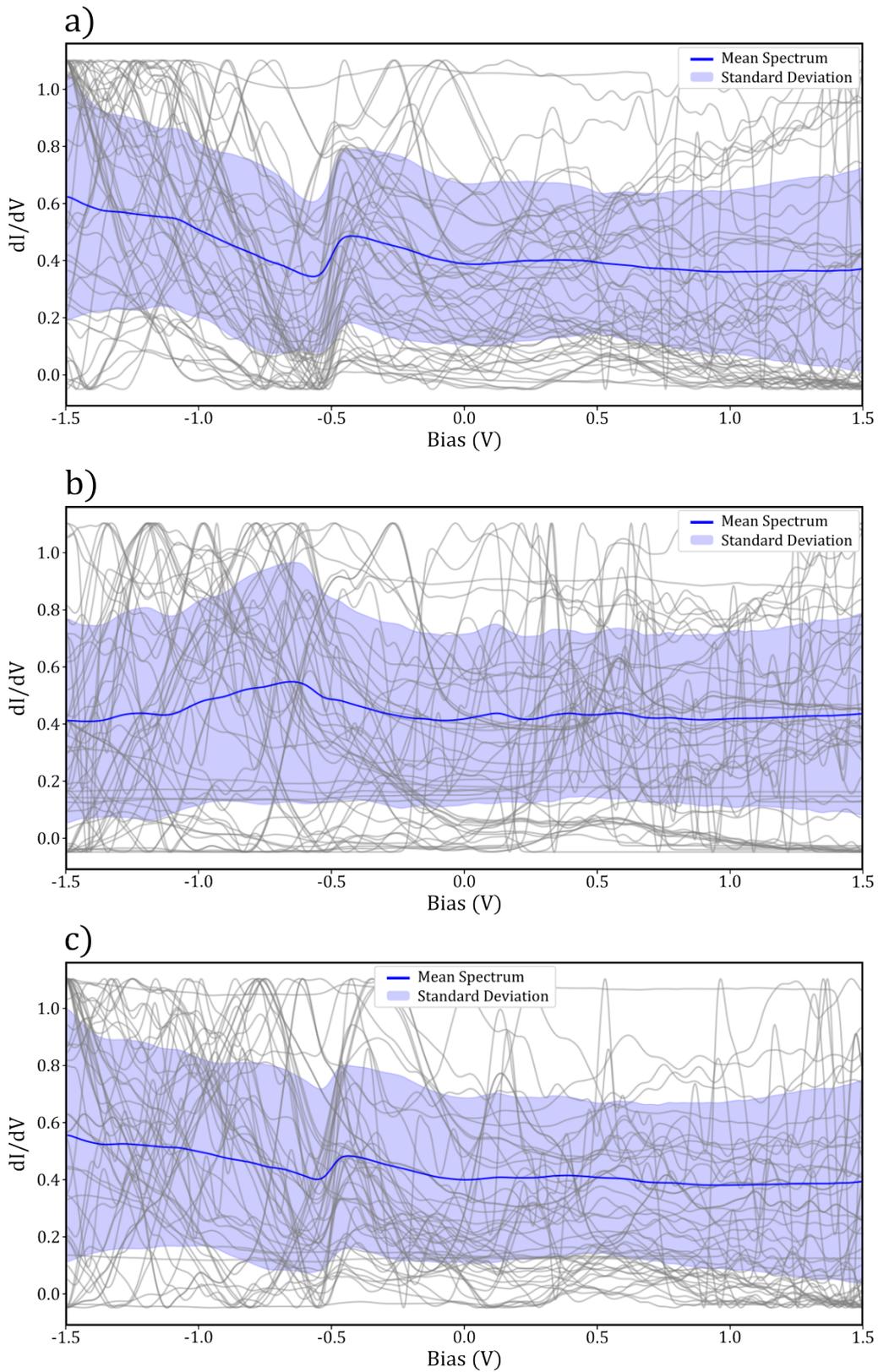


Figure 4: Samples of 50 normalised spectra taken over the clean Au(111) surface, mean (blue line) and standard deviation (shaded area) for a) Surface state peak visible, b) surface state not visible and c) binary “bad” labels.

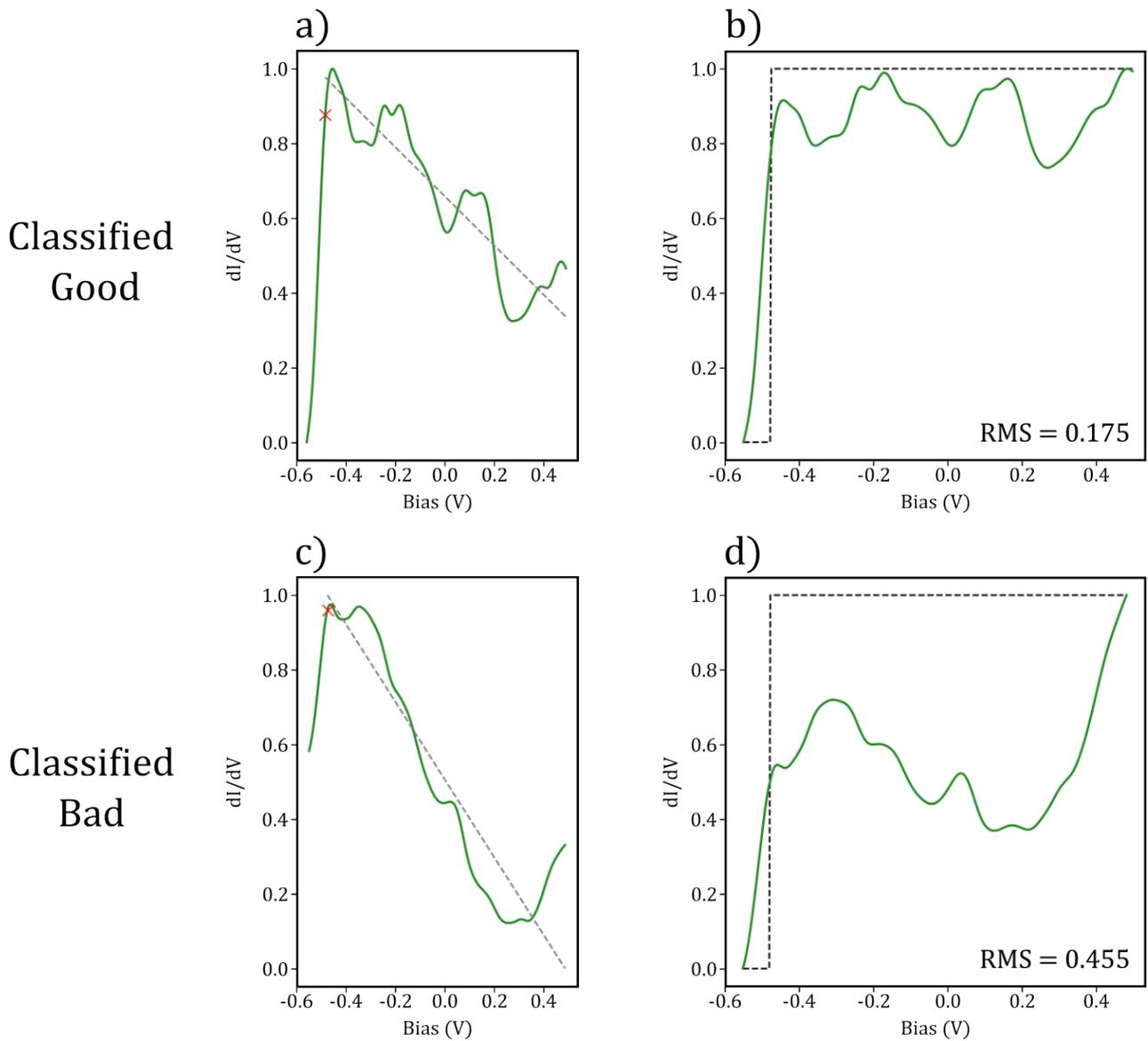


Figure 5: a) and b) show the categorisation window on a “good” and “bad” spectra respectively. The red crosses show the automatically located turning point of the step, and the dashed gray lines show the linear fit found past the step. b) and d) show the spectra in a) and b) with their respective linear fits subtracted. The black dashed curves show the ideal surface state step function.

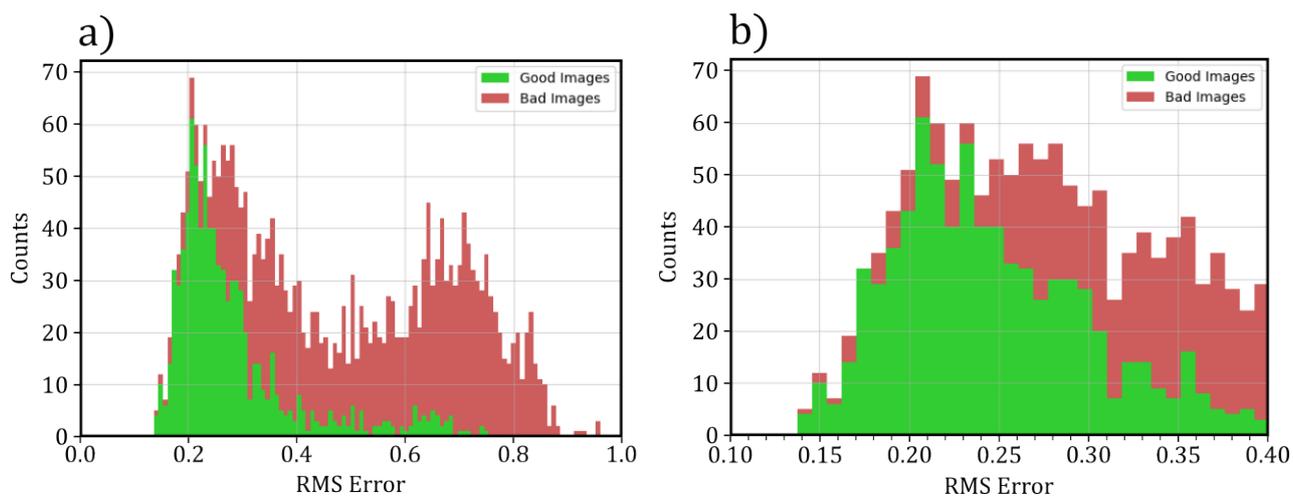


Figure 6: Stacked histogram made from labelled spectra, calculating the RMS error between each processed spectrum and an ideal step function. a) shows the full range of RMS, with b) showing the values between 0.1 and 0.4.

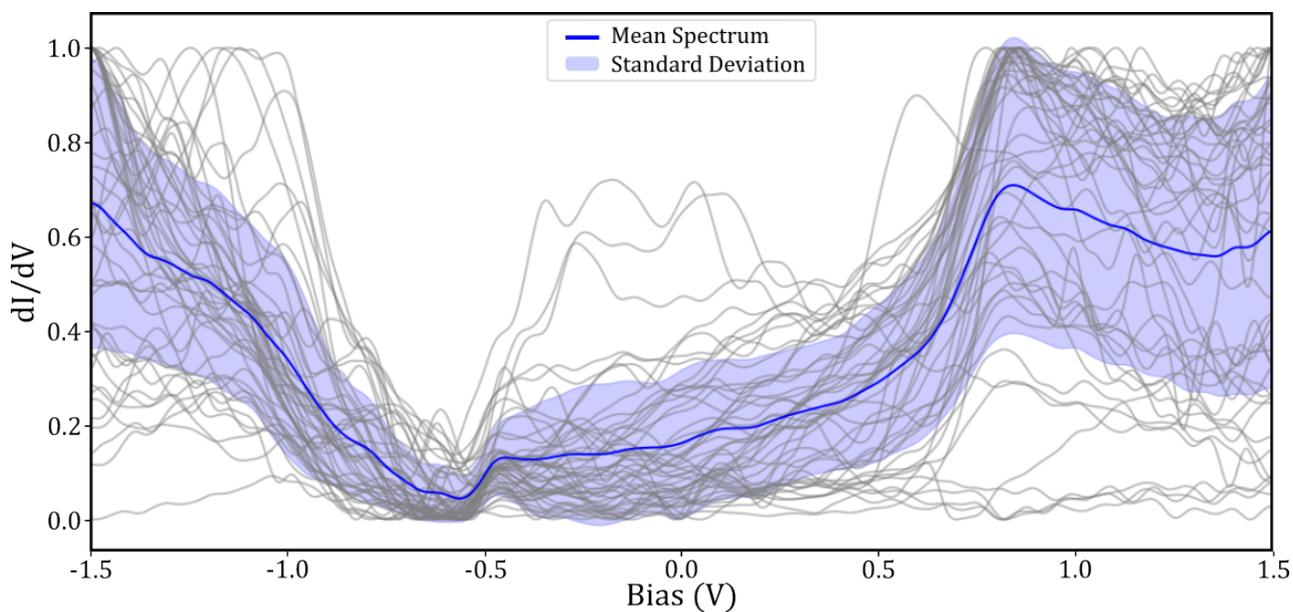


Figure 7: Gray curves show 49 normalised STS measurements taken over the centre of SnDown molecules taken with a “good” tip. Blue curve shows the mean and the standard deviation is shown in shaded purple.

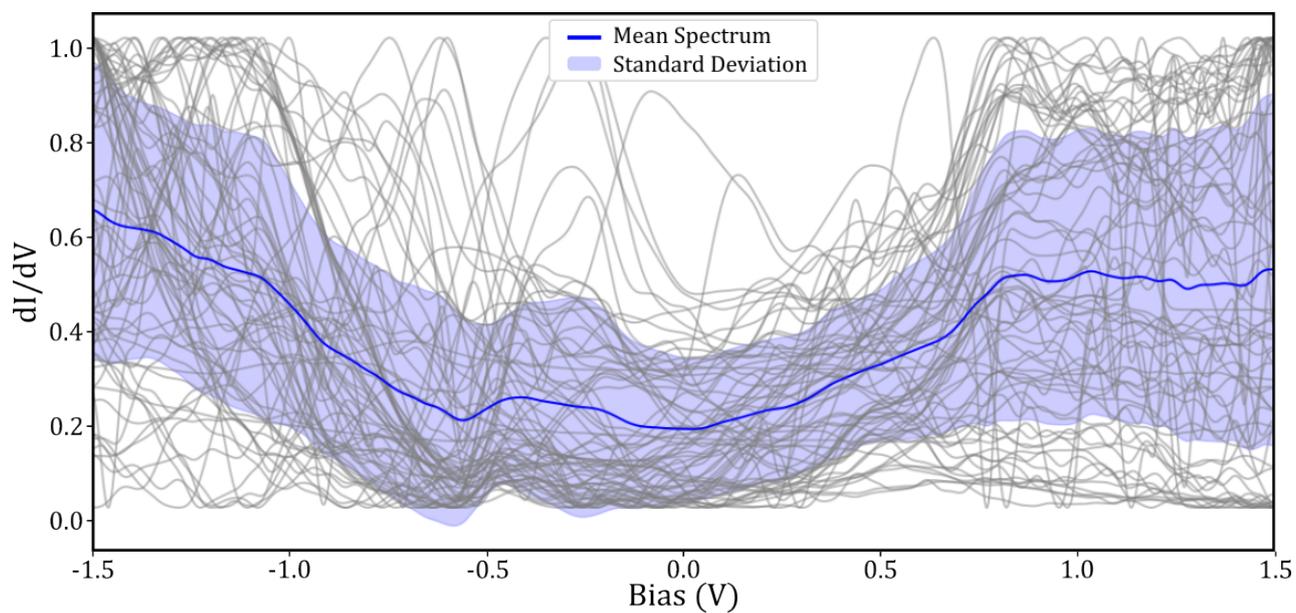


Figure 8: Gray curves show 71 normalised STS measurements taken over the centre of SnDown molecules taken with a “bad” tip. Blue curve shows the mean and the standard deviation is shown in shaded purple.

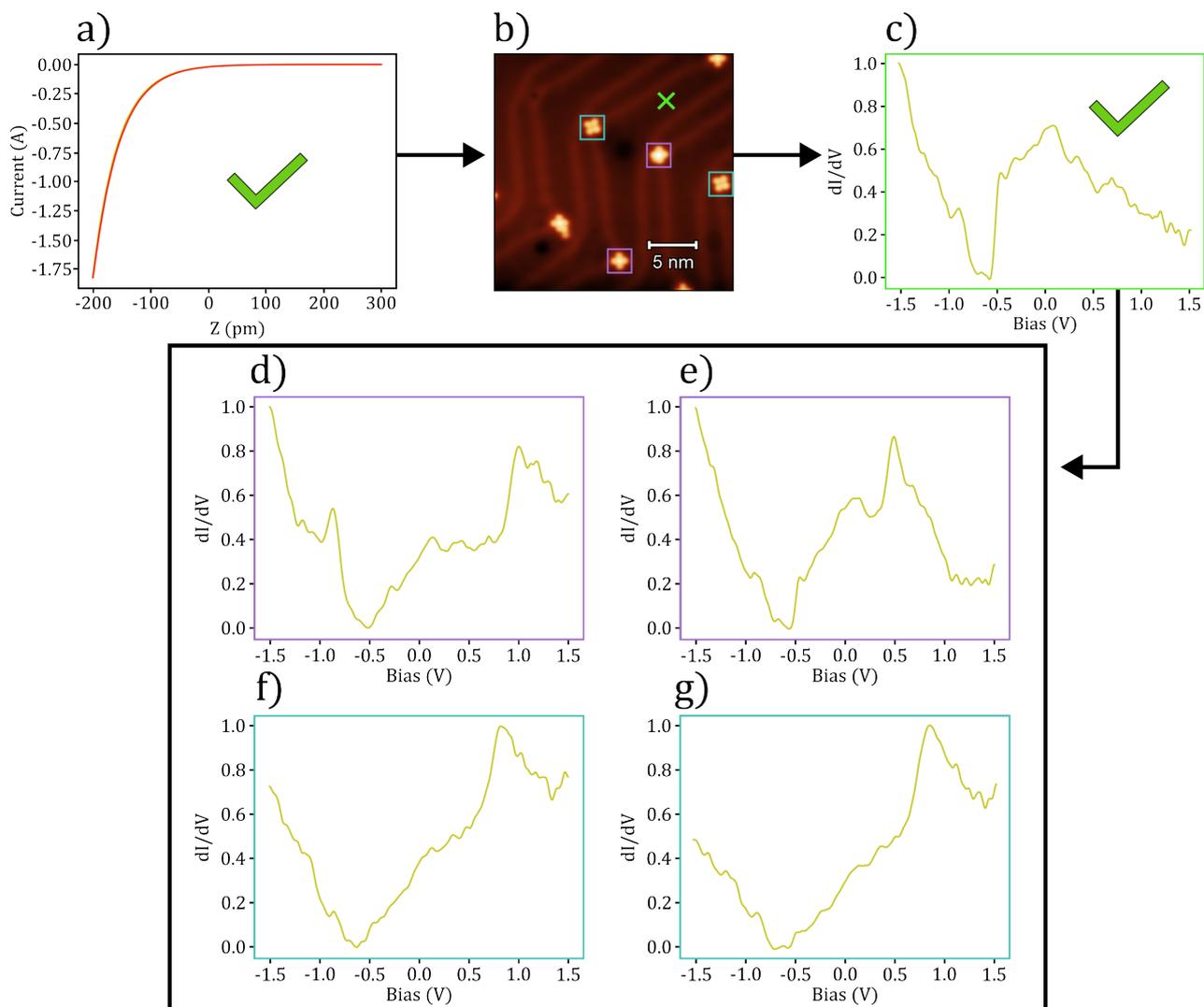


Figure 9: Example flow of an automated spectroscopy experiment taken over various SnPc molecules on the Au(111) surface. a) Initial $I(z)$ measurement is taken, where an exponential dependence is observed and so moves onto imaging, b). The tip is then classified to be “good” based on imaging, and so a clean area of the substrate is located (marked by a green cross), where a surface STS measurement is taken, c). This is then classified to be “good”, at which point the various orientations of SnPc are located (SnUp in pink boxes and SnDown in blue boxes), where STS measurements are taken as shown in d)-g). d)-e) correspond to measurements taken over SnUp molecules, while f)-g) correspond to SnDown. The script would then change the tip and repeat the steps, over different areas, varying the tip after each set of STS measurements (formed through *in situ* tip preparation).

References

- 379 1. Rizzo, D. J.; Veber, G.; Cao, T.; Bronner, C.; Chen, T.; Zhao, F.; Rodriguez, H.; Louie, S. G.;
380 Crommie, M. F.; Fischer, F. R. *Nature* 2018 560:7717 **2018**, 560, 204–208. doi:10.1038/
381 s41586-018-0376-8.
382
- 383 2. Gross, L.; Moll, N.; Mohn, F.; Curioni, A.; Meyer, G.; Hanke, F.; Persson, M. *PRL* **2011**, 107,
384 86101. doi:10.1103/PhysRevLett.107.086101.
- 385 3. Repp, J.; Meyer, G.; Stojković, S. M.; Gourdon, A.; Joachim, C. *Physical Review Letters* **2005**,
386 94, 026803. doi:10.1103/PHYSREVLETT.94.026803.
- 387 4. Repp, J.; Meyer, G.; Paavilainen, S.; Olsson, F. E.; Persson, M. *Science (New York, N.Y.)* **2006**,
388 312, 1196–1199. doi:10.1126/SCIENCE.1126073.
- 389 5. Bartels, L.; Meyer, G.; Rieder, K. H. *Applied Physics Letters* **1997**, 71, 213–215. doi:10.1063/
390 1.119503.
- 391 6. Lagoute, J.; Kanisawa, K.; Fölsch, S. *PHYSICAL REVIEW B* **2004**, 70, 245415. doi:10.1103/
392 PhysRevB.70.245415.
- 393 7. Liljeroth, P.; Repp, J.; Meyer, G. *Science (New York, N.Y.)* **2007**, 317, 1203–1206. doi:10.
394 1126/SCIENCE.1144366.
- 395 8. Tsukada, M.; Kobayashi, K.; Isshiki, N. *Surface Science* **1991**, 242, 12–17. doi:10.1016/
396 0039-6028(91)90234-J.
- 397 9. Kwapiński, T.; Jałochowski, M. *Surface Science* **2010**, 604, 1752–1756. doi:10.1016/J.SUSC.
398 2010.06.026.
- 399 10. Passoni, M.; Donati, F.; Bassi, A. L.; Casari, C. S.; Bottani, C. E. *Physical Review B - Con-*
400 *densed Matter and Materials Physics* **2009**, 79, 045404. doi:10.1103/PHYSREVB.79.045404.
- 401 11. Garcia, N.; Binh, V. T.; Purcell, S. *Surface Science Letters* **1993**, 293, L884–L886. doi:10.
402 1016/0167-2584(93)90226-9.

- 403 12. Palotás, K.; Hofer, W. A.; Szunyogh, L. *Physical Review B - Condensed Matter and Materials*
404 *Physics* **2011**, *83*, 214410. doi:10.1103/PHYSREVB.83.214410/FIGURES/4/MEDIUM.
- 405 13. Andreev, T.; Barke, I.; Hövel, H. *Physical Review B - Condensed Matter and Materials*
406 *Physics* **2004**, *70*, 205426. doi:10.1103/PHYSREVB.70.205426.
- 407 14. Thomas, J. C.; Rossi, A.; Smalley, D.; Francaviglia, L.; Yu, Z.; Zhang, T.; Kumari, S.; Robin-
408 son, J. A.; Terrones, M.; Ishigami, M.; Rotenberg, E.; Barnard, E. S.; Raja, A.; Wong, E.;
409 Ogletree, D. F.; Noack, M. M.; Weber-Bargioni, A. *npj Computational Materials* **2022** *8:1*
410 **2022**, *8*, 1–7. doi:10.1038/s41524-022-00777-9.
- 411 15. Wang, S.; Zhu, J.; Blackwell, R.; Fischer, F. R. *Journal of Physical Chemistry A* **2021**, *125*,
412 1384–1390. doi:10.1021/acs.jpca.0c10731.
- 413 16. Rashidi, M.; Wolkow, R. A. *ACS Nano* **2018**, *12*, 5185–5189. doi:10.1021/acsnano.8b02208.
- 414 17. Gordon, O.; D’Hondt, P.; Knijff, L.; Freeney, S.; Junqueira, F.; Moriarty, P.; Swart, I. *Review*
415 *of Scientific Instruments* **2019**, *90*, 103704. doi:10.1063/1.5099590.
- 416 18. Krull, A.; Hirsch, P.; Rother, C.; Schiffrin, A.; Krull, C. *Communications Physics* **2020**, *3*,
417 1–8. doi:10.1038/s42005-020-0317-3.
- 418 19. Barker, D. S.; Blowey, P. J.; Brown, T.; Sweetman, A. *ACS Nano* **2024**, *18*, 2384–2394. doi:10.
419 1021/ACSNANO.3C10597.
- 420 20. Yongfeng, W.; Kröger, J.; Berndt, R.; Hofer, W. A. *Journal of the American Chemical Society*
421 **2009**, *131*, 3639–3643. doi:10.1021/JA807876C.
- 422 21. Toader, M.; Hietschold, M. *Journal of Physical Chemistry C* **2011**, *115*, 3099–3105. doi:10.
423 1021/JP111478V.